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# Final Report Development of a Portable Infrared Emission Spectrometer AA460 PIRES

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# 1. Project Summary

The purpose of this Phase II research effort was to design and build a prototype field portable thermal infrared emission spectrometer (PIRES). This phase of work would build on the design concepts and analysis completed during Phase I.

The AA460 PIRES System is a field-portable thermal emission spectral radiometer which operates over the 2.5-14.5  $\mu$ m wavelength range. Its primary function is to characterize the spectral emission signatures of objects at reasonable terrestrial temperatures (0°-50°C). However, the instrument design is flexible enough to allow useful data collection to occur from objects outside this temperature range or from sources other than thermal radiators. While ideally suited for two-man operation in the field, the instrument can be transported and operated effectively by a single user. The AA460 PIRES System gives the user the capability to collect and display processed thermal emission data on-site. This eliminates the need for the user/researcher to remove a sample from its natural setting in order to accommodate laboratory analysis.

The AA460 PIRES instrument was fabricated, assembled, tested, and documented during this Phase II work period. Software development included a user friendly menu structure, control and communications processes. Testing included optical and spectral characterization of the instrument and general performance. The test results concluded that the instrument met most of the original design goals set forward in the Phase I work including size, weight, portability, rugged, menu driven operation and in-field analysis.

The area in which the instrument did not meet our design goals is in the signal-tonoise performance. Testing concluded that this poor SNR performance was due primarily
to data collection scheme and our inability to properly amplify, and stabilize the detector
signals. Modification to this collection scheme could improve the SNR performance and
verification would be provided by additional laboratory testing. This performance improvement verification is necessary before implementation of a redesign to the optical
head and its present data collection scheme.

# 2. Instrument Specifications

# 2.1. General Description

The AA460 PIRES System is a battery powered field-portable thermal emission spectral radiometer which operates over the 2.5-14.5  $\mu$ m wavelength range. The complete instrument consists of two main assemblies, the sensor head and the backpack. The complete operational system is shown in Figure 1.

The sensor head assembly, outlined in Figure 2, consists of the imaging optics, a 2-element sandwich-type detector, an all-attitude liquid nitrogen (LN<sub>2</sub>) dewar, an optical sight, a remote keyboard/display, two blackbody reference sources, instrument control electronics and mechanics, and a continuous variable filter (CVF) wheel. The backpack assembly primarily consists of a 12-volt battery and a portable control, processing, display, and storage computer. The backpack structure itself serves as a transportation and storage container for both the sensor head and the backpack assemblies. The backpack and the sensor head are electrically coupled by way of a multiconductor electrical cable over which numerous electrical signals travel.

# 2.1.1. Imaging Optics

The complete optical system consists of nine germanium lens elements and two deviating mirrors. These optical elements are used to:

- Define the instruments field-of-view (FOV)
- Image the energy through the CVF
- Image the energy onto the detector.

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Figure 1. AA460 PIRES System

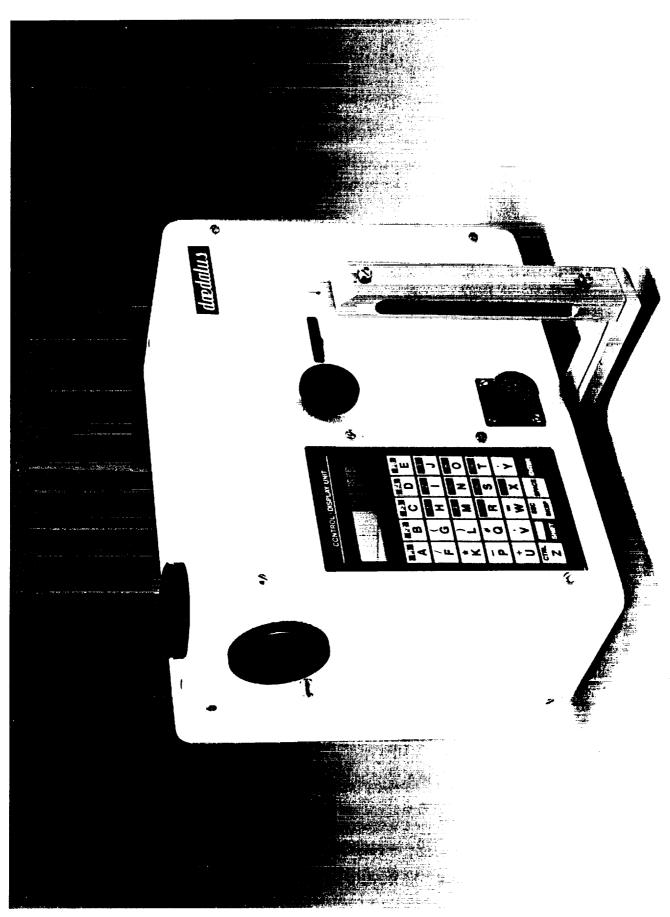


Figure 2. AA460 PIRES Sensor Head Assembly

The AA460 PIRES System provides the user with two selectable FOV's, 2° and 8°. The user can select to operate the instrument in either the 2° or 8° mode by throwing a mechanical switch located on the side of the sensor head.

# 2.1.2. Two-Element Detector

The detector which is used in the AA460 PIRES System is a 2-element sandwich-type detector. It consists of a thin layer of photovoltaic Indium Antimonide (InSb) over a layer of photoconductive Mercury Cadmium Telluride (MCT). InSb responds well in the 1-5.5  $\mu$ m wavelength range while MCT responds well in the 5.5-15  $\mu$ m wavelength range. Therefore, the use of a sandwich-type detector as described above ensures continuous satisfactory operation across the instruments 2.5-14.5  $\mu$ m wavelength range. Both InSb and MCT require cryogenic cooling to LN2 temperatures (77°K) for proper operation.

# 2.1.3. All-Attitude LN<sub>2</sub> Dewar

Integral to the AA460 sensor head is an all-attitude LN<sub>2</sub> dewar. This dewar houses the instruments InSb/MCT sandwich detector. Once filled with LN<sub>2</sub>, the dewar is specially designed to contain the LN<sub>2</sub> in all dewar positions, including upside down. This is a very important feature given the unpredictable nature of field operation. The all-attitude feature of the dewar makes transporting the instrument from site to site an easy procedure. In addition, it permits data collection to be performed in any conceivable instrument orientation. The hold time for a single charge of LN<sub>2</sub> is greater than 4 hours. That is, without refilling, a single charge of LN<sub>2</sub> will keep the instrument's sandwich detector at 77°K for 4 hours or more.

# 2.1.4. Optical Sight

A FOV-defining optical sight is contained in the sensor head. It is co-aligned with the axis of the instruments imaging optics axis. By looking through the eyepiece of the

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sight, the user is simultaneously provided with both a 2° FOV and 8° FOV outline. The two FOV's are respectively defined by the small and large square targets visible in the sight.

The sight is continuously adjustable over a target distance ranging from 18 inches to  $\infty$ . Due to the nature in which the instruments primary optics defines the systems IFOV, the sight is most accurate in predicting true instrument FOV when the target is far from the instrument ( $\infty$ ). At closer distances the instruments true FOV is slightly larger than what is defined by the sight.

# 2.1.5. Remote Keyboard/Display

A small membrane keyboard and 16-character by 2-line display is integral to AA460 PIRES sensor head. This remote keyboard/display provides the user with the capability of collecting spectral data without the need to access the full capability, menu-driven operational software found in the backpack mounted portable computer. With its lack of graphing capabilities and its limited ability to prompt, the remote keyboard/display provides the user access to only a subset of the total PIRES operational capabilities. However, the operational functions which are accessible from it are very useful in allowing the user to collect data when it is either inconvenient or impossible to use the backpack mounted computer.

# 2.1.6. Blackbody References

The AA460 sensor head contains two temperature monitored blackbody reference sources. Both of these blackbody reference sources are viewed by the instrument during the course of a data collection sequence. One of the references is an ambient blackbody reference source. It is neither actively heated or cooled. The other reference source is a hot blackbody reference. During data collection, this blackbody is heated above the ambient temperature to a control temperature. It is held at this control temperature

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during the entire data collection procedure. Ambient and hot blackbody spectral information obtained during data collection is used to calibrate the measured scene data.

The inclusion of two internal blackbody reference sources into the AA460 PIRES head eliminates the need to transport external blackbody references sources to the field site.

## 2.1.7. Continuous Variable Filter

Spectral dispersion in the AA460 PIRES system is achieved with the use of a continuous variable filter (CVF) wheel assembly. Manufactured out of three individual filter segments, the filter wheel assembly used in the PIRES sensor head provides for narrowband spectral selection across the  $2.5\text{-}14.5~\mu\mathrm{m}$  wavelength range. CVF's consist of optical substrates coated with multi-layer dielectric films. The spectral characteristics of the CVF is dependent on the thickness of the individual film layers and the refractive indices of both the film layers and the substrate materials. In the case of a CVF, the substrate/dielectric films combination produces a narrowband spectral filter which maintains a linear relationship between wavelength of peak transmission and angular position on the filter segment. The variable filtering action produced as the CVF assembly is rotated in the sensor head is used to isolate and measure scene spectral information.

# **2.1.8.** Battery

The AA460 PIRES system operates off a 12-volt battery. A Model PS-1265 12-volt sealed lead acid battery manufactured by Power Sonic Corp. is provided as part of the standard AA460 equipment. The battery is housed on the backpack structure. In addition, the system is mechanically and electrically designed to accept a 12-volt silver zinc battery. Silver zinc batteries are capable of delivering more data collection power as compared to lead acid batteries. However, they are significantly heavier and more expensive than lead acid type batteries. A battery design consisting of 8 Whitaker-

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Yardney Power Systems silver zinc cells No. LR20DC-5 can be supplied to the user as an option.

# 2.1.9. On-Board Computer

Integral to the AA460 backpack assembly is small, lightweight, low-power portable computer. This computer is used to perform several instrument functions including sensor head control, data processing, data storage, and data display. The backpack computer is the primary user interface to the PIRES system operational software. It consists of a 640 x 400 fixed LCD graphics display, a 3-1/2 inch floppy drive and a full function keyboard. Using the keyboard and display simultaneously, the user has access to the AA460 menu driven operational software. In addition to controlling the instrument, the user can also use the backpack computer to graph raw detector data as well as processed spectral information. Data collected by the instrument is stored on 1.4 Mbyte 3-1/2 inch floppy diskettes. In addition to executing PIRES specific control and processing software, the AA460 backpack computer is capable of running various user defined application programs. This enhances the instruments flexibility and usefulness in the field.

# 2.1.10. Backpack Structure

The backpack frame design of the AA460 PIRES system makes site-to-site transportation of the instrument a relatively easy procedure. The backpack computer and battery are permanently fastened to a lightweight fiberglass epoxied aluminum honeycomb material. Despite its light weight, this material is extremely strong and durable and, thus, it provides structural integrity to the overall system packaging scheme. During transportation the sensor head is also secured to this honeycomb composite.

In order to facilitate single user transportation of the complete system, the main PIRES structural member is then mounted to a standard backpack frame. To accommo-

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date a wide variety of instrument deployments, both in and out of the field, the backpack frame is designed to be quickly and easily removed from the rest of the unit. It is equally as easy to re-attach the pack frame to the rest of the backpack structure whenever needed.

# 2.2. Detailed Specifications

Listed below are some of the significant AA460 PIRES specifications.

Size: Sensor Head: 6.5 in. x 8.4 in. x 10.8 in.

Backpack: 9.0 in. x 16.4 in. x 33.5 in.

Weight: Sensor Head: 10.2 lbs.

Backpack: 25.0 lbs.

Environment: Operating Temperature:  $+5^{\circ}\text{C} - +45^{\circ}\text{C}$ 

Storage Temperature: -20°C - +60°C Humidity (non-condensing): 10% - 90%

Spectral Coverage:  $2.5 - 14.5 \mu m$ 

Spectral Resolution:  $\leq 2\%$  of  $\lambda$ . Given a single line source at  $\lambda_0$ , the half-

power bandwidth of the displayed spectrum would be

 $\leq (0.02) \lambda_o$ .

Noise Equivalent Temper-

ature Difference (NETD):

 $\leq 0.2^{\circ}$ C fom 4-14  $\mu$ m (proposed)

Field of View: User selectable, 2° or 8° square.

Data Acquisition Time: Data Collection: Variable, depending on number of

samples averaged.

Battery Life:  $\geq 200$  spectra with fully charged lead acid battery.

≥ 1000 spectra with fully charged silver zinc battery

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Liquid Nitrogen Hold

Time:

≥ 4 hours

Dynamic Range:

12 bits

Data Display:

LCD graphics: 640 x 400 pixels

Screen Size: 9 in. x 6 in.

Cable Lengths:

10 ft. (standard)

25 ft. (optional)

Data Storage:

1.4 Mbyte, 3-1/2 in. floppy diskette

Data Transfer:

Removable 3-1/2 in. floppy diskette; parallel printer

port available for hardcopy transfer of LCD graphics

information.

Sensor Head Mounting:

Hand-held or tripod mounting.

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# 3. Performance

# 3.1. Testing

Testing and alignment of various system components were accomplished during fabrication. These included the following:

- Alignment of the complete imaging optical system
- CVF positioning for optimum spectral bandwidth
- Spectral calibration.

These tests showed that the design had achieved the two and eight degree fields of view and that the spectral bandwidth was approximately 2 percent of the wavelength and for approximately  $2.60 \,\mu\text{m} - 14.5 \,\mu\text{m}$  with no gaps in coverage.

The spectral calibration was performed using a scanning McPherson monchrometer. The field of view of the instrument was filled with energy from the monchrometer using an off-axis parabola. Three separate gratings for the monchrometer were required to cover the broad spectral range. Data for wavelength position verse encoder step was determined for the  $2.55\mu m - 12.5 \mu m$  by rotating the CVF and viewing the response on an oscilloscope. At  $12.5 \mu m$  we could no longer get enough energy from the monchrometer and extrapolated the encoder position versus wavelength for the rest of the third segment. During this calibration process, spectral bandwidth was determined at several points on the CVF by scanning the monchrometer and determining the half-power points. At 9.0  $\mu m$  the measured bandwidth was 1.96% and at 6.0 and  $2.6 \mu m$  in the bandwidth was 1.98% and 1.52% respectively.

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# 3.2. Signal/Noise Performance

The instrument signal-to-noise performance was calculated using a utility program embedded into the PIRES software. This program is used to convert raw data collected into wavelength dependent NETD information. The user selects the number of data samples to be collected and the NETD program statistically computes the measured sensitivity performance of the system. This raw data is then processed off line. Due to the extensive amounts of processing associated with this program, it is recommended that it be executed on either a 286 or 386 AT computer equipped with a math co-processor.

Using the data file produced during SNR data acquisition, we will statistically compute an SNR value at each wavelength for each detector. As mentioned before, these SNR algorithms will be contained in a stand-alone computer program. In order to statistically determine SNR in this manner, the signal energy which is supplied to the system must remain constant throughout the course of the SNR data acquisition operation. In our case, the signal of interest is the DIFFERENCE in radiance between the hot and ambient blackbody. The only way this signal can remain constant is for BOTH the blackbody temperatures to remain constant. It is anticipated that both the hot and ambient blackbody temperatures will vary somewhat during SNR data acquisition. The first step then in performing a SNR calculation is to go through the thermistor data line by line and select those data lines wherein both the hot and ambient blackbody counts fall within some user defined range. Hopefully the selected data will comprise a large subset of the total data collected. After this "good" data has been identified, the following operations will be performed on it.

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$$A_{\lambda_1}n = \left(InSb\ Hot\ \right) - \left(InSb\ Ambient\ \right);$$

$$B_{\lambda_1\gamma}n = \left(MCTHot\right) - \left(MCTAmbient\right)\ mean\ InSb\ signal\ at\ \lambda = U_{\lambda} = \frac{\sum\limits_{n}A_{\lambda_1}n}{\#\ of\ good\ InSb\ data\ lines\ at\ \lambda}$$

$$mean\ MCT signal\ at\ \lambda = \nu_{\lambda} = \frac{\sum\limits_{n} B_{\lambda_{1}n}}{\#\ of\ good\ MCT\ data\ lines\ at}$$

$$standard\ deviation\ of\ InSb\ a\ \lambda = \sigma_{v\lambda} = \sqrt{\frac{\sum\limits_{n} + \left[A\lambda_{1}n - U\lambda\right]^{2}}{\#\ of\ good\ InSb\ data\ lines}}$$

standard deviation of MCT at 
$$\lambda = \sigma_{v_{\lambda}} = \sqrt{\frac{\sum\limits_{n} \left[B_{\lambda_{1}n} - v_{\lambda}\right]^{2}}{\# of good MCT data lines}}$$

Now,

$$SNR\lambda$$
,  $InSb = \frac{U\lambda}{\sigma_{tt\lambda}}$ 

$$SNR_{\lambda}$$
, $MCT = \frac{\gamma_{\lambda}}{\sigma_{\gamma_{\lambda}}}$ 

These SNR values are outputted in tabular format listed below in Table 1.

There are 335 good wavelengths

TABLE 1.

Detector Detector Detector Detector Detector System 1 1 Signal 2 Signal 1 Noise 2 Noise 1 NETD 2 NETD NETD	lengt
1 Signal 2 Signal 1 Noise 2 Noise 1 Nois 3 Nois	
0.000 41.105 0.000 1207.042 3.400e+038 2.320e+002 3.400e+038	
0.000 431.250 255.660 1837.423 3.400e+038 3.366e+001 3.400e+038	
0.000 0.000 402.200 1864.269 3.400e+038 3.400e+038 3.400e+038	2.643
0.000 59.529 380.687 2346.890 3.400e+038 3.114e+002 3.400e+038	2.661
0.000 329.500 453.697 1401.702 3.400e+038 3.361e+001 3.400e+038	
0.000 184.167 365.321 1683.803 3.400e+038 7.223e+001 3.400e+038	
0.000 0.000 351.940 2052.802 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 356.845 2291.100 3.400e+038 3.400e+038 3.400e+038	
0.000 186.188 460.532 1168.768 3.400e+038 4.959e+001 3.400e+038	
0.000 0.000 404.451 2254.334 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 238.748 2531.687 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 396.056 2348.850 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 359.523 1919.415 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 396.920 1862.990 3.400e+038 3.400e+038 3.400e+038	
	2.855
0.000 0.000 313.250 2320.989 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 324.097 2350.664 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 336.544 3110.237 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 294.510 2093.049 3.400e+038 3.400e+038 3.400e+038	•
0.000 0.000 297.243 2737.141 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 414.212 2838.410 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 341.295 2575.936 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 291.413 2220.848 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 325.479 2400.281 3.400e+038 3.400e+038 3.400e+038	
	3.032
	3.049
	3.067
0.000 0.000 375.985 2348.316 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 395.616 2151.186 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 369.860 1668.381 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 296.843 3329.350 3.400e+038 3.400e+038 3.400e+038 0.000 0.000 377.304 3048.591 3.400e+038 3.400e+038 3.400e+038	3.155
·	
0.000 0.000 391.691 2657.062 3.400e+038 3.400e+038 3.400e+038 0.000 0.000 319.557 2745.442 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 319.337 2743.442 3.400e+038 3.400e+038 3.400e+038 0.000 0.000 306.768 2709.475 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 300.768 2709.475 3.400e+038 3.400e+0400	
0.000 0.000 415.450 3178.424 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 415.430 3176.424 3.400e+038 3.400e+038 3.400e+038 3	
0.000 0.000 299.471 3096.990 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 489.495 3265.723 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 418.111 2845.475 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 409.657 3277.231 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 399.243 2772.811 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 358.369 2946.914 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 385.409 3091.740 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 356.598 2815.750 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 396.776 2766.342 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 354.330 3131.015 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 461.624 2484.955 3.400e+038 3.400e+038 3.400e+038	
0.000 0.000 348.971 2683.976 3.400e+038 3.400e+038 3.400e+038	

```
3.400e+038 3.509
                                3040.640 3.400e+038 3.400e+038
              0.000
                      263.065
   0.000
                                                                             3.526
                                                     3.400e+038
                                                                 3.400e+038
                                         3.400e+038
                                3005.191
                      303.765
              0.000
   0.000
                                                                 3.400e+038
                                                                             3.544
                                                     3.400e+038
                                         3.400e+038
                                3724.349
              0.000
                      319.242
   0.000
                                                                 3.400e+038
                                                                            3.562
                                                     3.400e+038
                                         3.400e+038
                      422.971
                                3429.390
              0.000
   0.000
                                                                             3.579
                                                                 3.400e+038
                                                     3.400e+038
                                3084.913
                                         3.400e+038
                      329.686
   0.000
              0.000
                                                                             3.597
                                                     3.400e+038
                                                                 3.400e+038
                                         3.400e+038
                                2794.133
                      448.368
              0.000
   0.000
                                                                 3.400e+038
                                                                             3.615
                                                     3.400e+038
                                         3.400e+038
                      350.915
                                2502.283
              0.000
   0.000
                                                                             3.632
                                                     3.400e+038
                                                                 3.400e+038
                                          3.400e+038
                      322.107
                                3140.209
   0.000
              0.000
                                                                 3.400e+038
                                                                             3.650
                                         3.400e+038
                                                     3.400e+038
                      441.574
                                3564.958
              0.000
   0.000
                                                                             3.668
                                                     3.400e+038
                                                                 3.400e+038
                                3318.100 3.400e+038
              0.000
                      360.545
   0.000
                                                                 3.400e+038
                                                                             3.685
                                         3.400e+038
                                                     3.400e+038
                      383.318
                                3165.013
   0.000
              0.000
                                                                 3.400e+038
                                                                             3.703
                                                     3.400e+038
              0.000
                                          3.400e+038
                      452.859
                                2750.702
   0.000
                                                                             3.721
                                                                 3.400e+038
                                          3.400e+038
                                                     3.400e+038
                                3006.905
                      419.844
   0.000
              0.000
                                                                             3.738
                                                     3.400e+038
                                                                 3.400e+038
                                         3.400e+038
                      392.247
                                3186.423
              0.000
   0.000
                                                     3.400e+038
                                                                 3.400e+038
                                                                             3.756
                                         3.400e+038
                                3593.315
                      379.302
   0.000
              0.000
                                                     3.400e+038
                                                                 3.400e+038
                                                                             3.774
                                          3.400e+038
              0.000
                      396.429
                                2902.375
   0.000
                                                                 3.400e+038
                                                                             3.791
                                3171.403 3.400e+038
                                                     3.400e+038
                      359.109
   0.000
              0.000
                                                     3.400e+038
                                                                 3.400e+038
                                                                             3.809
                                3249.781 3.400e+038
                      465.979
              0.000
   0.000
                                                                             3.827
                                                     3.400e+038 3.400e+038
                                         3.400e+038
                      368.784
                                2290.432
   0.000
              0.000
                                                                 3.400e+038
                                                                             3.844
                                                     3.400e+038
                                3230.048
                                          3.400e+038
                      393.776
   0.000
              0.000
                                                                             3.862
                                         3.400e+038
                                                     3.400e+038
                                                                 3.400e+038
              0.000
                      397.220
                                3118.079
   0.000
                                                                 3.400e+038
                                                                             3.880
                                                     3.400e+038
                                3320.266 3.400e+038
                      389.814
              0.000
   0.000
                                                                             3.897
                                                                 3.400e+038
                                         3.400e+038
                                                     3.400e+038
              0.000
                      361.482
                                3658.056
   0.000
                                                                             3.915
                                                                 3.400e+038
                                3480.255 3.400e+038
                                                     3.400e+038
   0.000
                      404.946
              0.000
                                                                             3.933
                                                     3.400e+038
                                                                 3.400e+038
                                3729.156 3.400e+038
              0.000
                      487.707
   0.000
                                                                 3.400e+038
                                                                             3.950
                                                     3.400e+038
                                3097.784
                                          3.400e+038
                      366.024
   0.000
              0.000
                                                                             3.968
                                3195.125
                                          3.400e+038
                                                     3.400e+038
                                                                 3.400e+038
                      437.745
              0.000
   0.000
                                3199.342 3.400e+038 3.400e+038
                                                                 3.400e+038
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   0.000
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                              1924.225 3.400e+038 1.149e+001 1.149e+001 12.77
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                              1488.435 7.317e+001 9.362e+000 9.362e+000 12.83
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                              2000.597 2.568e+001 1.338e+001 1.338e+001 13.01
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                              1776.698 1.125e+002 8.215e+000 8.215e+000 13.07
 33.857
         1708.476
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                              2166.021 2.957e+002 1.373e+001 1.373e+001 13.12
                     431.363
 11.524
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                                       3.400e+038 1.228e+003 1.228e+003 13.18
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          896.476
                     425.334
 68.905
                              2570.092 3.400e+038 1.548e+001 1.548e+001 13.30
                     496.333
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                              2283.327 3.400e+038 2.632e+001 2.632e+001 13.36
          685.375
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                              1996.302 3.400e+038 4.664e+001 4.664e+001 13.41
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          254.278
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                              2201.471 2.511e+002 2.089e+001 2.089e+001 13.53
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                     493.305
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                              1824.976 2.389e+002 8.833e+001 8.833e+001 13.94
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                              2053.823 9.373e+001 3.400e+038 3.400e+038 14.17
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                              1545.721 1.850e+002 3.400e+038 3.400e+038 14.23
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                              1580.433 4.674e+001 3.400e+038 3.400e+038 14.29
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                              1583.658 2.363e+001 7.678e+001 7.678e+001 14.35
155.700
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                              2066.026 1.494e+001 4.562e+001 4.562e+001 14.41
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                              1874.919 4.301e+001 2.238e+001 2.238e+001 14.46
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 63.800
            0.000
                     532.582
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As you can see from this data, the instrument has measurable performance in only two wavelength areas. These are the  $4.4 \,\mu\text{m} - 5.5 \,\mu\text{m}$  and  $8.2 - 11.5 \,\mu\text{m}$  region which correspond to peak response areas of the InSb and MCT detector elements. The original instrument performance analysis completed in the Phase I work predicated NEDT performance of approximately one order of magnitude better than the measured values.

The Phase I analysis was based on a slightly different instrument design that chopped the input radiation against the reference sources and CVF motion was controlled by a stepper motor. This design required an electronic bandwidth of approximately 400 Hz. The Phase II development work concluded that with CVF stepper motion it would be impossible to implement a data collection time of 3 seconds. Tests showed that a minimum of 6 seconds might be possible with a highly developed mechanical damping mechanism needed to provide reliable non-jittering steps of the CVF. A more reasonable design would provide data acquisition in about 10 seconds. We felt that this would be unacceptable for a hand-held instrument. In addition, the dual choppers required to provide two blackbody reference sources had made the overall optical head size very large and heavy for a hand-held instrument. An alternate concept was pursued during the Phase II work and a design developed which eliminated the chopping of the input radiation against references. This scheme used a small reference and target mirrors to view first the ambient reference for a complete CVF rotation, then the target for another CVF rotation and finally the hot reference for a final CVF rotation.

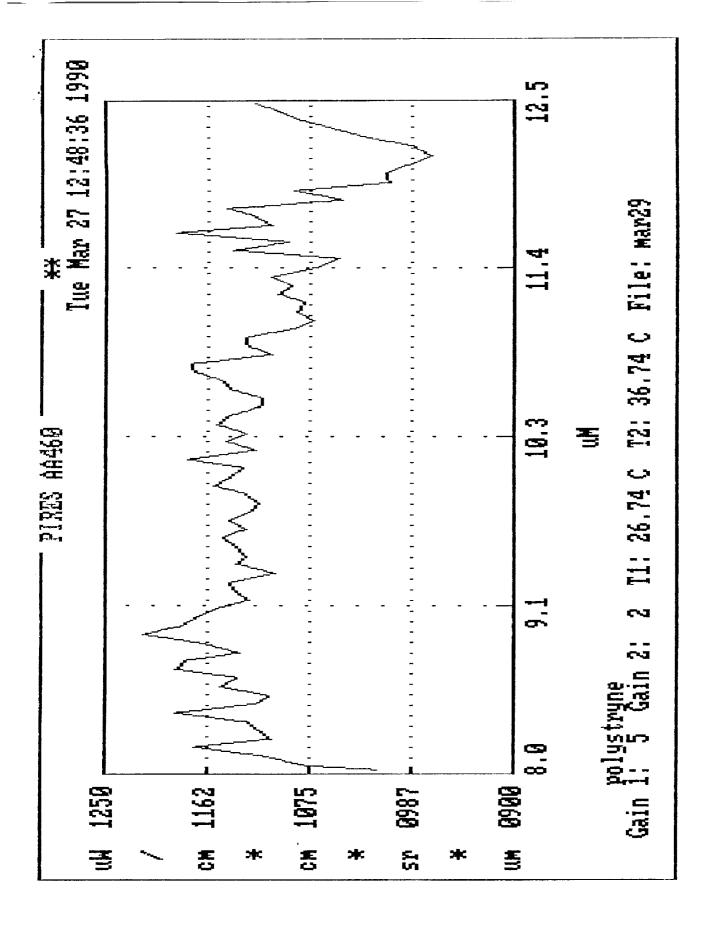
This data acquisition scheme required that the CVF be rotated at a high rate of approximately 1000 revolutions per minute. This increased CVF rotation rate causes the system bandwidth requirements to increase. We concluded that we could average multiple CVF samples to maintain the original SNR requirements. Several consequences of this data collection scheme is the large variation in analog detector signals from segment to segment of the CVF. This is mainly due to the variation in detector response over the large wavelength region. The second problem area is the stabilization of the detector

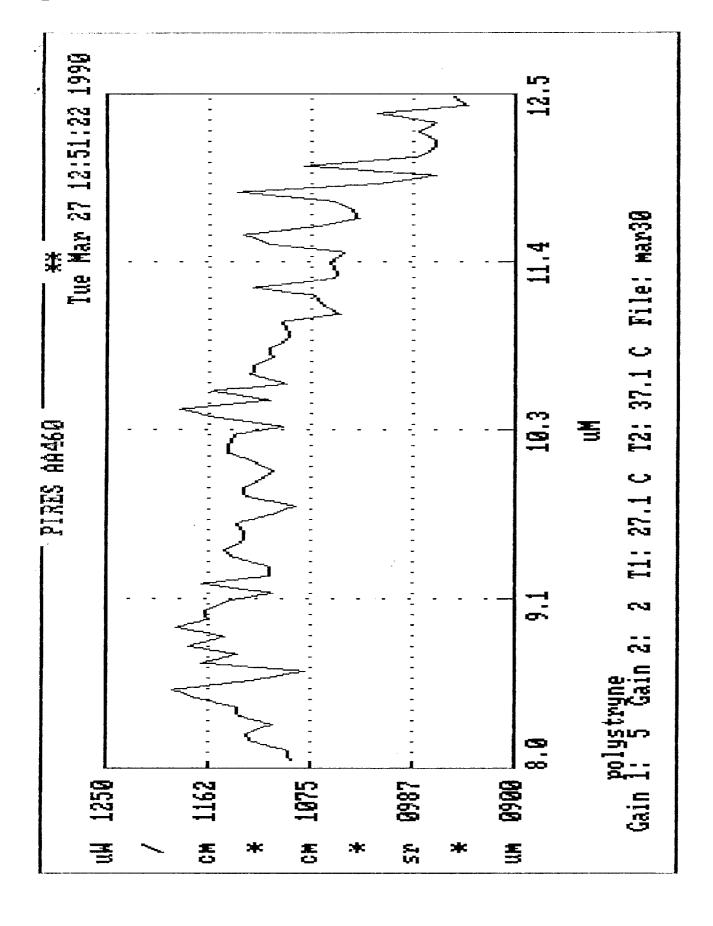
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signals as the data is being averaged over all three targets. The analog signal processing include an analog clamp to stabilize the detector signals. This clamp was applied three times per CVF revolution when the detectors were looking at surface mirror section installed in between the CVF segments. These cold spikes are very constant and stable except for detector/preamp white noise and system noise. The large analog detector variation and cold spike sampling stabilization limited us to relatively low preamp gains to keep the analog signal out of saturation, including the cold spike clamping region. Our inability to properly amplify the detector signals and maintain the detector stabilization while data is gathered over the references and target is the main reason for poor instrument performance. The original data collection scheme of optical choppers and CVF steppers would allow proper amplification of these very low level analog signals without the need for clamping and offsetting circuits. In addition, a lock-in amplifier signal processing design could be implemented where in only the chopped frequency of the analog signal is used for processing.

As a further verification of poor instrument performance, spectra of polystyrene were obtained using the PIRES instrument; these are included in the following graphs. As you can see from these graphs, the processed radiance spectral data appears noisy and inconsistent for data taken over the same sample.

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# 4. Recommendations

The PIRES instrument, in its present form, needs improvement in its SNR performance in all wavelength regions. This further instrument development could be approached in the following ways.

The first approach is to continue with the present data collection scheme and improve SNR by better detector stabilization electronics for many rotations of the CVF, reducing systematic instrument noise in the analog detector signals, and providing better analog signal processing including improved offset removal and amplification of detector signal. This approach I feel will only lead to slight SNR improvements in the present instrument.

A better alternative would be to redesign the fore optics of the instrument to include optical choppers as originally proposed. This altered data collection scheme would then allow a dramatic decrease in system bandwidth requirements and, hence, improved SNR. In addition, I would recommend the CVF motor be changed to start and stop (stepper motor driven). This could easily be accomplished by replacement of the brushless DC motor with a stepper type. The present belt-driven driver pulleys could be modified to provide proper CVF motion per motor step. Signal-to-noise improvement using this approach should be verified by temporarily modifying the PIRES instrument to incorporate a stepper CVF drive and an ambient blackbody chopper inserted in the primary optical path. Detector signals could be bandwidth limited using laboratory filter and RMS-to-DC converted. This converted data could then be digitized and analyzed to verify improved system performance before undertaking the major optical head redesign to incorporate these features.

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